



Subaru Coronagraphic Extreme Adaptive Optics

すばるコロナグラフ極限補償光学装置

Exoplanet Imaging and Characterization with Subaru SCExAO & TMT PSI



Presenter: Olivier Guyon (guyon@naoj.org)

NAOJ & ABC team:

Thayne Currie, Jun Hashimoto, Takayuki
Kotani, Tomoyuki Kudo, Masayuki Kuzuhara,
Julien Lozi, Toshiyuki Mizuki, Naoshi
Murakami, Jun Nishikawa, Motohide
Tamura, Taichi Uyama

SCExAO and PSI teams



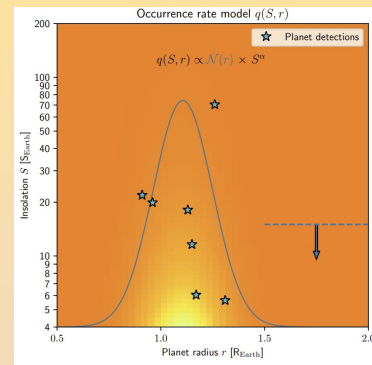
Credit: NAOJ



- Transit and RV surveys have revealed that a **significant fraction (~10-50%) of stars host potentially habitable planets**
- Terrestrial planets appear especially abundant around M-type stars
- Several nearby stars have potentially habitable planets - See Ultra-Doppler presentation

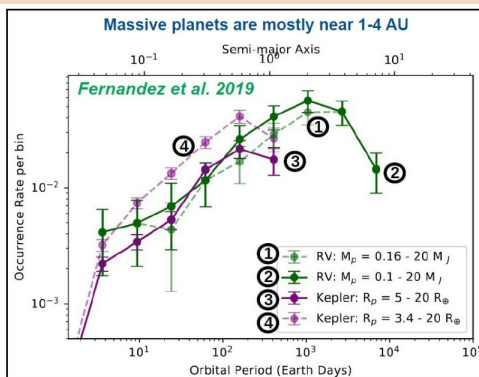
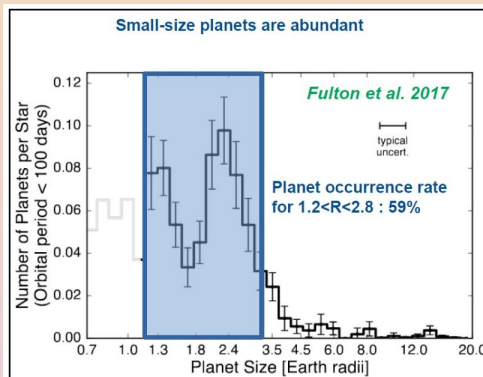
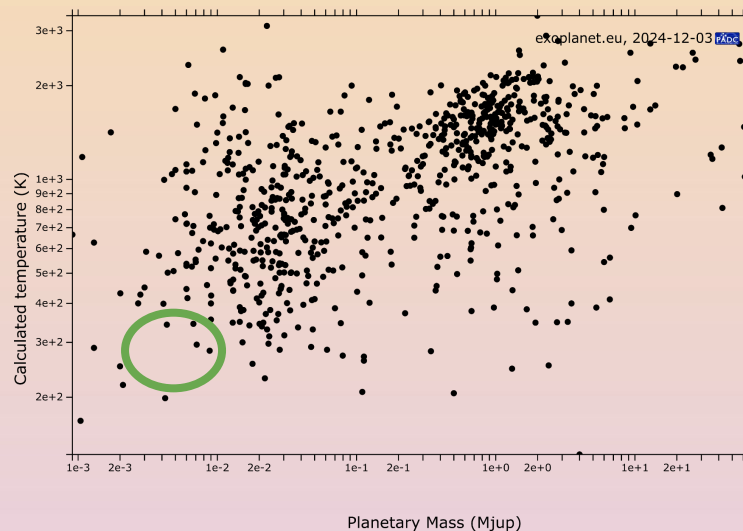
Current limitations:

- Most (>90%) nearby planets do not transit
- RV does not provide atmospheric characterization
- Direct imaging (currently) limited to young massive exoplanets

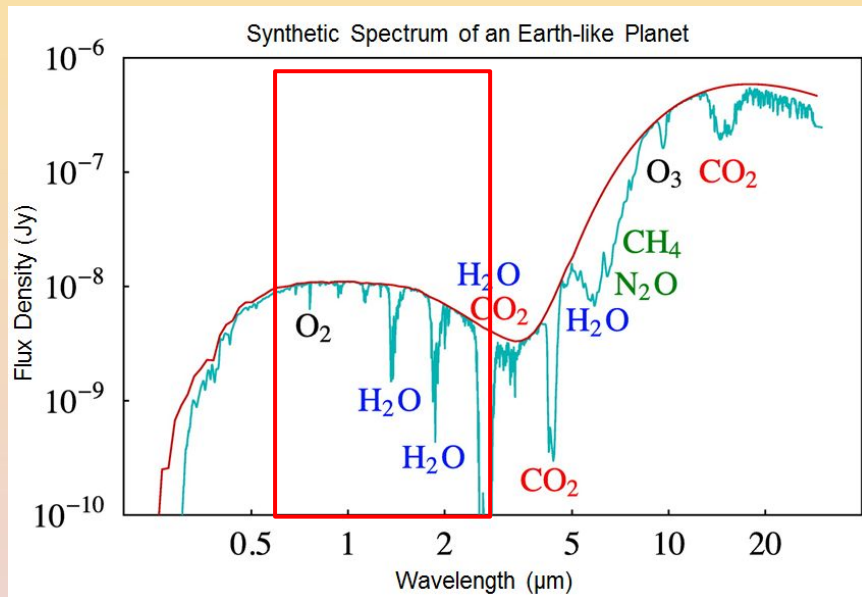


Terrestrial exoplanets population informed from TESS

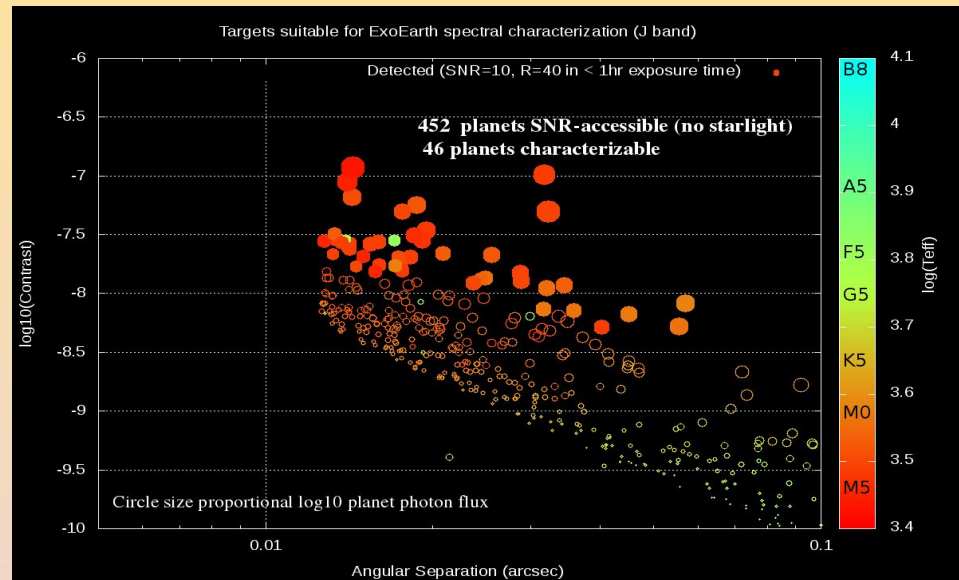
Ment & Charbonneau 2024



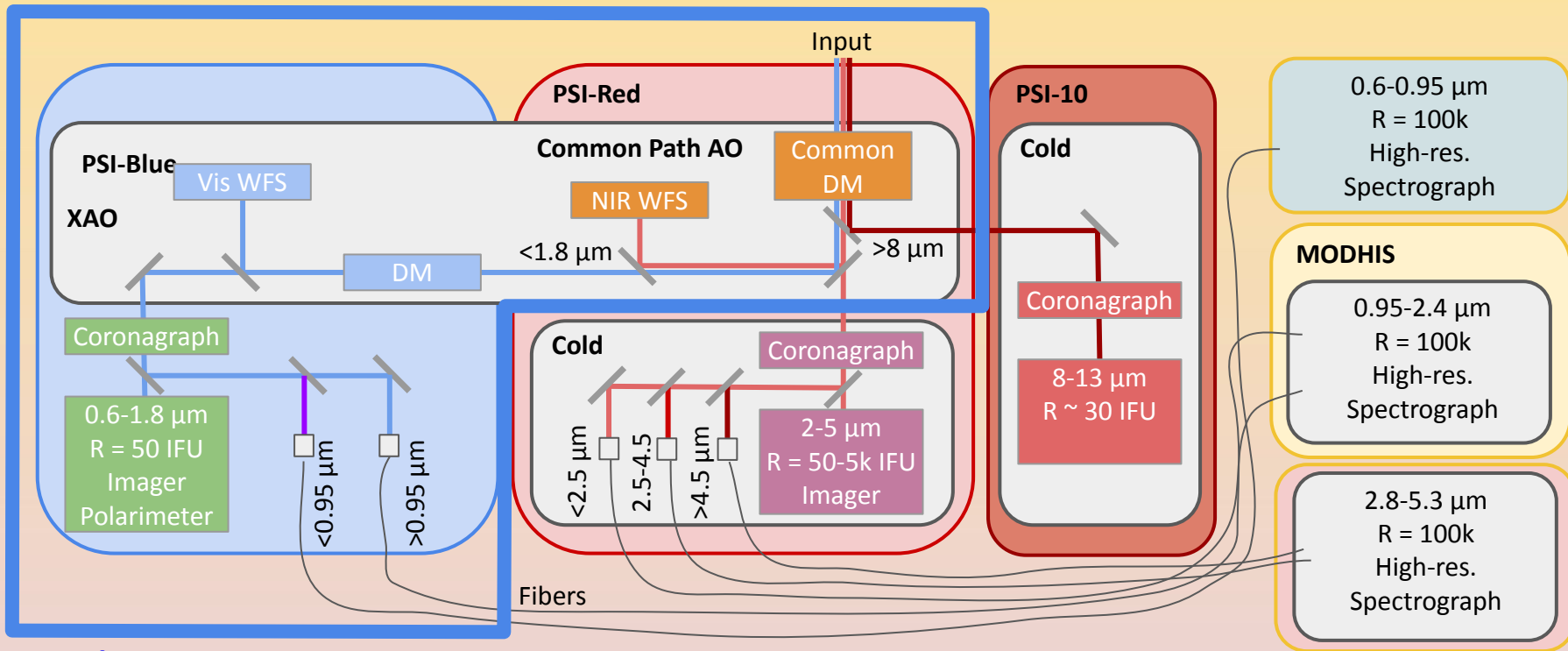
Spectroscopic characterization in VIS/NIR



Atmosphere characterization & detection of biomarkers: H_2O , O_2 , CH_4 , CO_2



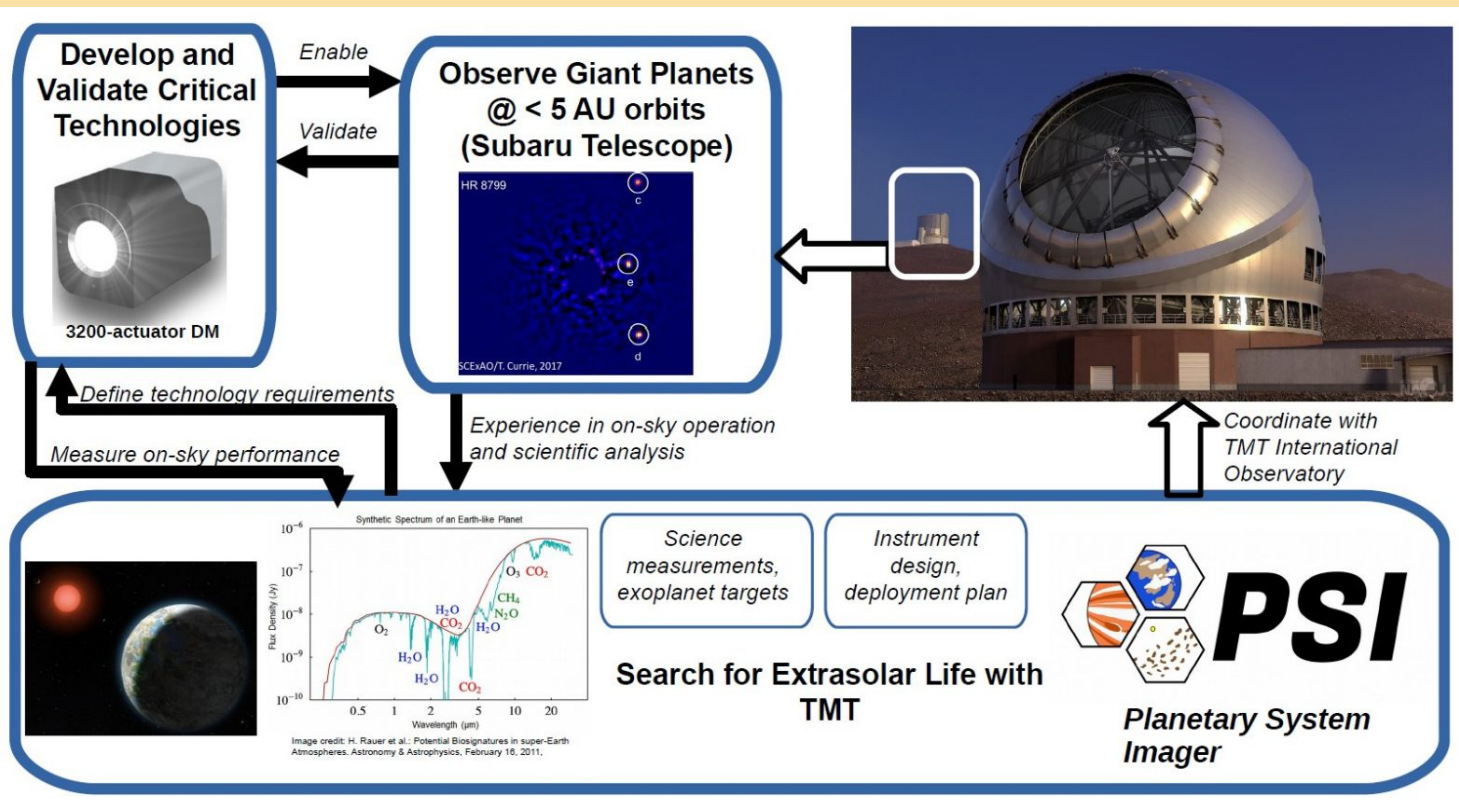
ExAO system on 30-m telescope can reach habitable planet characterization sensitivity (SNR=10 @ R=40 in 1hr) around 46 nearby stars.



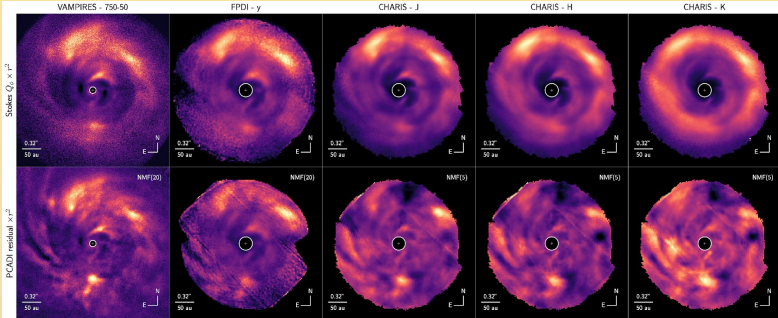
NAOJ/ABC lead PSI-BLUE - prime science goal: habitable exoplanet imaging & spectroscopy

US/Canada leads PSI-RED

Credit: PSI team



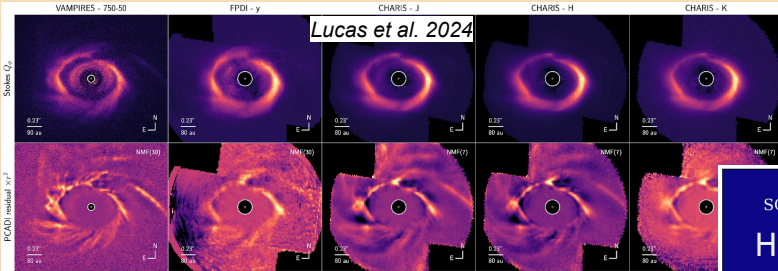
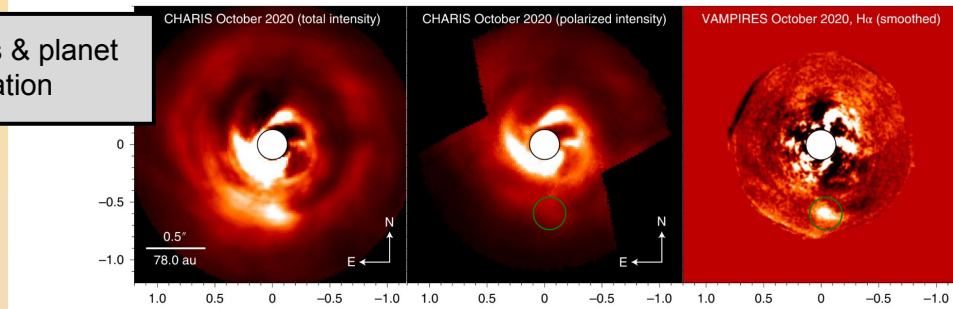
We will work closely with the existing Planetary Systems Imager (TMT-PSI) instrument team to develop the *PSI-BLUE* part of the instrument, optimized for habitable planet observations.



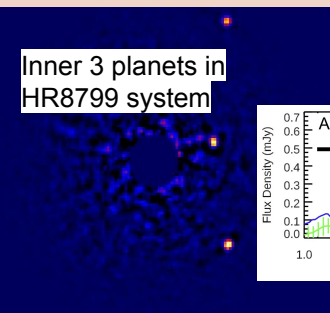
“Images of embedded Jovian planet formation at a wide separation around AB Aurigae”, *Currie et al. 2022*

NATURE ASTRONOMY

Disks & planet formation

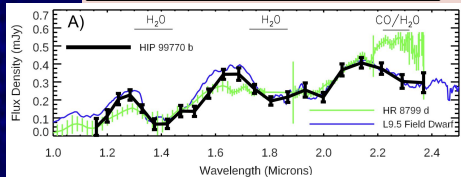


Lucas et al. 2024

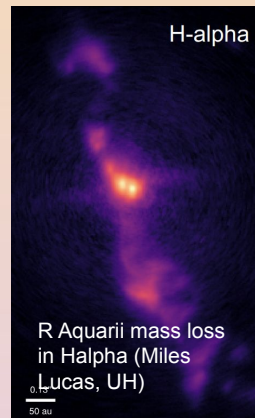
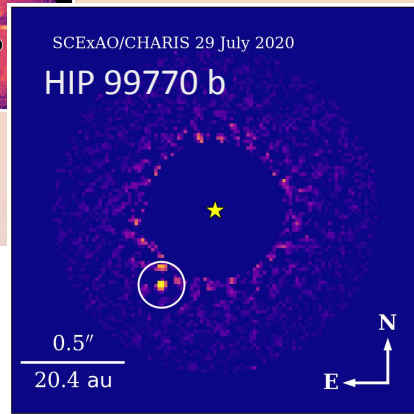


Inner 3 planets in HR8799 system

Exoplanet Imaging & Spectroscopy



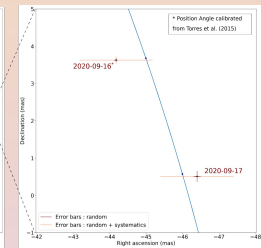
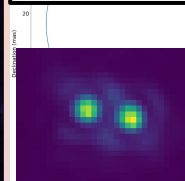
Currie et al. 2023, Science



R Aquarii mass loss in H-alpha (Miles Lucas, UH)

Stellar Physics

0.046" separation



“Single-aperture spectro-interferometry in the visible at the Subaru telescope with FIRST: First on-sky demonstration on Keho’oea (α Lyrae) and Hokulei (α Aurigae)”, *Vievard et al. 2023*

Develop and validate on-sky critical technologies and systems

@ Subaru:

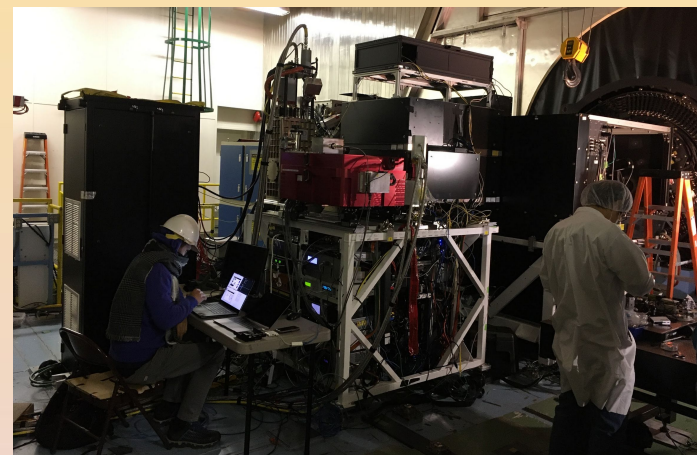
- High-performance AO correction
- Focal plane wavefront control
- Starlight suppression (coronagraphy, photonic)
- Fast low-noise detectors (CMOS, ImAPD, MKIDS)
- Signal processing
- Exoplanet Spectroscopy

-> Enables new science @ Subaru: mapping inner planetary systems (architecture, evolution, composition)

Existing SCEXAO system become prototype for PSI. Early career personnel training (science and technology) to prepare for PSI.

-> **Critical for balancing PSI performance, risk and schedule**

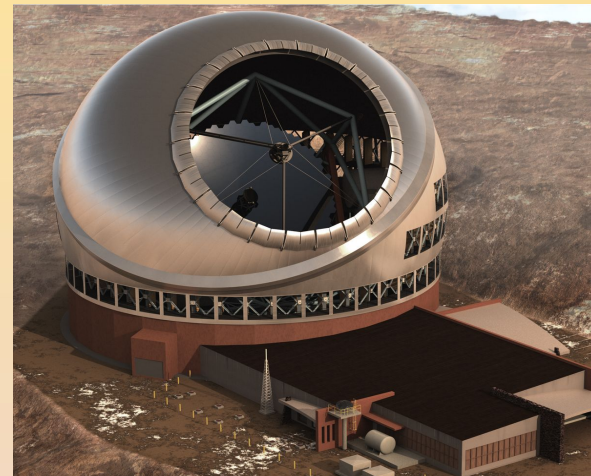
-> **Critical to grow the TMT-PSI Japan community**



Re-use SCEXAO experience, hardware and software to the maximum extent possible for TMT-PSI

Why NAOJ & ABC ?

- NAOJ and ABC are **leaders in high contrast imaging & spectroscopy**, leveraging SCExAO, IRD & Japan-based lab R&D
- **Strong exoplanet science research background** with ABC, post-SEEDS, ALMA
- Program needs engagement at the national (Japanese) and international (TMT partners) levels

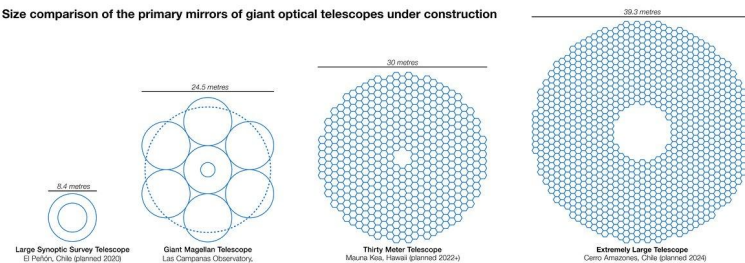


Credit: TMT International Observatory

Why TMT ? (vs. GMT or ELT)

- Maunakea is **better site** than Chilean sites with longer coherence time (no jet stream) and lower telluric absorption of key biomarkers (H₂O and O₂).
- TMT has **friendly pupil** with low segmentation and small central obstruction, and high-efficiency 3-mirror design

Size comparison of the primary mirrors of giant optical telescopes under construction



Credit: ESO

International Context:

- GMT planning for first-light instrument (GMagAO-X) with focused capabilities (2030+ ?)
- ESO developing PCS for ELT, currently in technology development phase (2040 ?)
- NASA developing HWO mission (2040+ ?) with complementary capabilities (will probe G-type stars, not M-type stars)

TMT-PSI will be the **only habitable exoplanet imaging instrument in the northern hemisphere**

TMT-PSI will likely come online after GMagAO-X, but will be **more capable** (better site, larger telescope, more ambitious instrument).

TMT-PSI can be deployed **before ELT's PCS**, leveraging NAOJ/ABC experience.

We are aiming at developing TMT-PSI for deployment soon after TMT first light (2034 ??).

2025-2030: SCEXAO-based R&D program in high contrast imaging/spectroscopy

2026-2030: TMT-PSI instrument design & component prototyping

2029-2033: TMT-PSI-blue build

2025-2030: SCEXAO-based R&D program in high contrast imaging/spectroscopy

USD ~\$2M per yr: Grants from Japan (JSPS), US (NSF, NASA, private), Australia, Europe & Canada. This is continuation of existing R&D program

2026-2030: TMT-PSI(-blue) instrument design & component prototyping

USD ~\$1M per yr: Grants from Japan & TMT-PSI members

2029-2033: TMT-PSI-blue build

USD ~\$25M total: Once TMT construction resumed, NAOJ+ABC submit MEXT request (Gaisan Yokyu)

NAOJ contributions:

- 2x FTEs (NAOJ faculty) to assist for SCEXAO science operation/use and R&D coordination, proposal preparation, Japan TMT-PSI-blue promotion, team building engaging young researchers. Also needed to help manage possible NAOJ contribution to NASA-led HWO mission.
- Operation support (SCEXAO: PI-type -> facility instrument ?)
- Support with software development, data management, computing infrastructure to support science & R&D, provide access & support to Japanese researchers

Before ELT era, SCEXAO will remain world-leading exoplanet imaging system with growing capabilities.

During TMT-PSI operation: TMT-PSI performance will ultimately be vastly superior to Subaru/SCEXAO for exoplanet imaging and spectroscopy, but TMT observing time will be in high demand, and TMT capabilities will take time to ramp up.

- Subaru/SCEXAO should run **survey(s) to identify exoplanets** (giant gas planets) and characterize them (spectroscopy) when possible.
- RV with Subaru/SCEXAO + Ultra-Doppler to **identify targets for TMT-PSI & HWO**. This is required to confirm directly imaged planets and **estimate exoplanet masses**.
- Subaru/SCEXAO can remain **competitive in short wavelength (optical/blue)**, for which AO becomes extremely challenging on a 30 m aperture.
- Subaru/SCEXAO can run **specialized modes** that may not be (initially?) deployed on TMT-PSI to keep the instrument simple: fast PDI (disks), interferometry.

AO3k + LTAO can be evolved to a visible-light MCAO system to provide diffraction-limited (15mas resolution) visible light performance over ~30" field of view.