

Selection of Amino Acid Homochirality in Stellar Environments

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Because they are asymmetric, amino acids (an example of which is shown in Figure 1) have two chiral states - or mirror images - referred to as “left-handed” and “right-handed” forms. With very few notable exceptions, nearly all life on earth utilizes only left-handed amino acids, a phenomenon referred to as biomolecular homochirality. Additionally, analyses of inclusions in meteoritic carbonaceous chondrites have revealed numerous amino acids. The discovery that *meteorites not only contain amino acids, but can have amino acids favoring left handed amino acids* is a possible indicator that amino acid chiral selection may have extra-terrestrial origins and has led to several hotly contested theories. Some of the amino acids so synthesized have an excess of a few percent of the left-handed chirality that is observed in earthly amino acids. It is generally accepted that if some mechanism can introduce an imbalance in the populations of the left- and right-handed forms of any amino acid, successive autocatalysis can amplify this enantiomeric excess to ultimately produce a single form. What is not well understood is the mechanism by which the initial imbalance is produced and the means by which it produces the left-handed chirality observed in the amino acids. *The origin of biomolecular homochirality is one of the top questions in science today* [1].

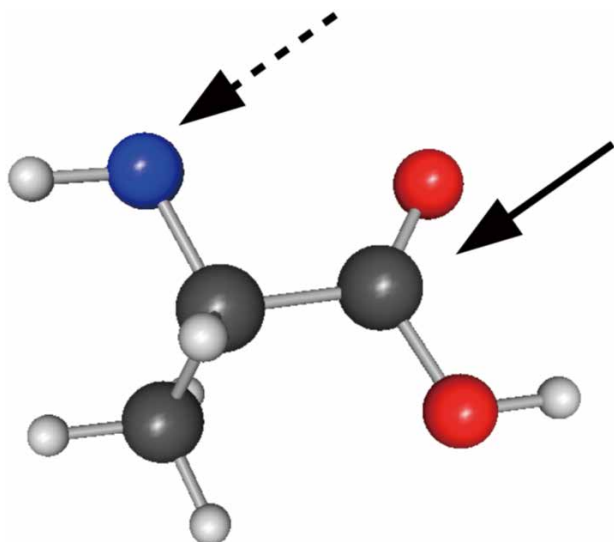


Figure 1: The left-handed enantiomer of the amino acid, alanine, L-alanine. The amine group is indicated by the dashed arrow and the carboxyl group by the solid arrow.

A model has been created which predicts the origins of amino acid homochirality in an emergent stellar phenomenon involving lepton interactions in magnetic fields [2]. Multiple sites capable of implementing this models have been proposed [3].

The results of an evaluation are shown in Figure 2 for the amino acid isovaline (in cationic form) for a scenario representing conditions of a meteoroid in the vicinity of a neutron star. The antineutrino reaction rates are given by the ratios $f = 0, 0, 10^{-9}, 5 \times 10^{-9}, 10^{-8}, 2 \times 10^{-8}$, and 5×10^{-8} , where f is defined as the ratio of the neutrino interaction rate to the nitrogen nuclear spin relaxation rate.

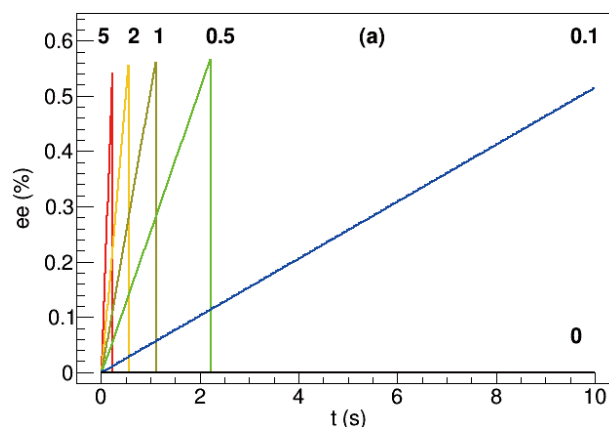


Figure 2: Enantiomeric excess of isovaline in a high-field, high flux environment for several different neutrino interaction rates.

In this figure, the relative amount by which one chiral state outnumbers another is given by the “enantiomeric excess,” ee , defined as $ee \equiv (N_L - N_D)/(N_L + N_D)$ where N_L and N_D represent the number of left- and right-handed amino acids, respectively, of a specific type in an ensemble. The ees computed in this model are similar to those found in meteorites. It is possible that amino acid ees have their origin in space, while subsequent autocatalysis via chemical or biological means can increase the ee to the homochiral values observed today.

References

- [1] Kennedy, D., Norman, C.: 2005, *Science*, **309**, 78.
- [2] Famiano, M. A., et al.: 2018, *Astrobiology*, **18**, 190.
- [3] Boyd, R. N., et al.: 2018, *ApJ*, **856**, 26.